The κ -Carroll particle from 3d gravity

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Outline:



A point particle coupled to 3d gravity

- 3d gravity as the Chern-Simons theory
- Gravitating particle in flat spacetime



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A point particle coupled to 3d gravity

- 3d gravity as the Chern-Simons theory
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- There is a variety of problems in the search for quantum gravity
- 3d gravity is a simpler version of ordinary general relativity
- It has no local degrees of freedom i.e. no gravitational waves
- Dynamics can be reintroduced via a nontrivial spacetime topology
- The 3d Newton's constant has the dimension of inverse mass
- Momentum space of a point particle coupled to 3d gravity is curved
- Curved momentum space is characteristic to some of the approaches to quantum gravity, especially "relative locality"

Chern-Simons formalism Particle in flat spacetime

De Sitter gauge group

For the cosmological constant $\Lambda>0$ the local isometry group of space-time is de Sitter group SO(3, 1). Its algebra has the commutators

$$\begin{bmatrix} J_{\mu}, J_{\nu} \end{bmatrix} = \epsilon_{\mu\nu\sigma} J^{\sigma} , \qquad \begin{bmatrix} J_{\mu}, P_{\nu} \end{bmatrix} = \epsilon_{\mu\nu\sigma} P^{\sigma} ,$$

$$\begin{bmatrix} P_{\mu}, P_{\nu} \end{bmatrix} = -\Lambda \epsilon_{\mu\nu\sigma} J^{\sigma} , \qquad (1)$$

where $\mu, \nu, \sigma = 0, 1, 2$. Introducing new generators $S_{\mu} \equiv P_{\mu} + \sqrt{\Lambda} \epsilon_{\mu 0\nu} J^{\nu}$ we may rewrite it as

$$[J_{\mu}, J_{\nu}] = \epsilon_{\mu\nu\sigma} J^{\sigma}, \qquad [J_{\mu}, S_{\nu}] = \epsilon_{\mu\nu\sigma} S^{\sigma} + \sqrt{\Lambda} (\eta_{\nu 0} J_{\mu} - \eta_{\mu\nu} J_{0}),$$

$$[S_{\mu}, S_{\nu}] = \sqrt{\Lambda} (\eta_{\mu 0} S_{\nu} - \eta_{\nu 0} S_{\mu}).$$
(2)

Thus group elements $\gamma \in SO(3, 1)$ can be locally factorized into

$$\gamma = \mathfrak{j}\mathfrak{s} = (\iota_3 + \iota^{\mu}J_{\mu})(\xi_3 + \xi^{\nu}S_{\nu}), \qquad (3)$$

with $j \in SO(2, 1)$, $\mathfrak{s} \in AN(2)$ and $\iota_3^2 + \frac{1}{4}\iota_{\mu}\iota^{\mu} = 1$, $\xi_3^2 - \frac{\Lambda}{4}\xi_0^2 = 1$. We also have the natural scalar product

$$\langle J_{\mu}S_{\nu}\rangle = \eta_{\mu\nu}, \qquad \langle J_{\mu}J_{\nu}\rangle = \langle S_{\mu}S_{\nu}\rangle = 0.$$
 (4)

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Chern-Simons action of 3d gravity

Instead of the metric $g_{\alpha\beta}$ gravity may be described by the vielbein $e_{\alpha}^{\ \mu}$ and spin connection $\omega_{\alpha}^{\ \mu\nu}$, defined through

$$\boldsymbol{e}_{\alpha}^{\ \mu}\boldsymbol{e}_{\beta}^{\ \nu}\eta_{\mu\nu} = \boldsymbol{g}_{\alpha\beta}\,,\qquad \omega_{\alpha}^{\ \mu\nu} = \boldsymbol{e}_{\beta}^{\ \mu}\partial_{\alpha}\boldsymbol{e}^{\beta\nu} + \boldsymbol{e}_{\beta}^{\ \mu}\Gamma_{\ \alpha\gamma}^{\beta}\boldsymbol{e}^{\gamma\nu}\,. \tag{5}$$

For 3d gravity we can introduce the gauge field, which is the Cartan connection

$$\boldsymbol{A} = \omega^{\mu} \boldsymbol{J}_{\mu} + \boldsymbol{e}^{\mu} \boldsymbol{P}_{\mu} \,, \tag{6}$$

where $e^{\mu} = e_{\alpha}^{\ \mu} dx^{\alpha}$ and $\omega^{\mu} = -\frac{1}{2} \epsilon^{\mu}_{\ \nu\sigma} \omega_{\alpha}^{\ \nu\sigma} dx^{\alpha}$, and the Einstein-Hilbert action can be written as the Chern-Simons gauge theory

$$S = \frac{k}{4\pi} \int \left(\langle dA \wedge A \rangle + \frac{1}{3} \langle A \wedge [A, A] \rangle \right) , \qquad (7)$$

with the coupling constant $k \equiv \frac{1}{4G}$.

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Coupling of a particle to 3d gravity

If we separate time \mathbb{R} and space S then A may be split into $A = A_t dt + A_S$. The action of gravity with a point particle is given by

$$S = \int dt \ L = \frac{k}{4\pi} \int dt \int_{S} \left\langle \dot{A}_{S} \wedge A_{S} \right\rangle - \int dt \left\langle C \ h^{-1} \dot{h} \right\rangle + \int dt \int_{S} \left\langle A_{t} \left(\frac{k}{2\pi} F_{S} - hC \ h^{-1} \delta^{2}(\vec{y}) \ dy^{1} \wedge dy^{2} \right) \right\rangle.$$
(8)

The algebra element $C = m J_0 + s S_0$ encodes the particle's mass m and spin s and the group element h describes the particle's motion. The spatial curvature $F_S = dA_S + [A_S, A_S]$ satisfies the constraint

$$\frac{k}{2\pi}F_{\mathcal{S}} = h\mathcal{C} h^{-1}\delta^2(\vec{y}) dy^1 \wedge dy^2, \qquad (9)$$

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i.e. vanishes everywhere except a singularity at the particle's worldline.

Chern-Simons formalism

Alekseev-Malkin construction

Space S may be decomposed into the disc D surrounding the particle, with coordinates $r \in [0, 1]$, $\phi \in [0, 2\pi]$, and the asymptotic empty region \mathcal{E} (for r > 1), with the common boundary Γ at r = 1. Solving the curvature constraint we find that on \mathcal{E} the connection has the form

$$A_{S}^{(\mathcal{E})} = \gamma d\gamma^{-1} , \qquad (10)$$

while on \mathcal{D} it is given by

$$A_{S}^{(\mathcal{D})} = \bar{\gamma} \, \frac{1}{k} \mathcal{C} d\phi \, \bar{\gamma}^{-1} + \bar{\gamma} d\bar{\gamma}^{-1} \,, \quad \bar{\gamma}(r=0) = h \,, \tag{11}$$

where γ , $\bar{\gamma}$ are some gauge group elements. The continuity of A_{S} across Γ , i.e. $A_{S}^{(\mathcal{D})}|_{\Gamma} = A_{S}^{(\mathcal{E})}|_{\Gamma}$ leads to the sewing condition

$$\gamma^{-1}|_{\Gamma} = N \, \boldsymbol{e}^{\frac{1}{k}\mathcal{C}\phi} \bar{\gamma}^{-1}|_{\Gamma} \,, \tag{12}$$

where N = N(t) is an arbitrary gauge group element.

Poincaré gauge group

In the limit $\Lambda \to 0$ the group $\mathrm{SO}(3,1)$ becomes the Poincaré group $\mathrm{ISO}(2,1) \simeq \mathfrak{so}(2,1)^* \rtimes \mathrm{SO}(2,1), \mathfrak{so}(2,1)^* \simeq \mathbb{R}^3$, with the algebra

$$[J_{\mu}, J_{\nu}] = \epsilon_{\mu\nu\sigma} J^{\sigma}, \qquad [J_{\mu}, P_{\nu}] = \epsilon_{\mu\nu\sigma} P^{\sigma}, \qquad [P_{\mu}, P_{\nu}] = 0.$$
(13)

Group elements $\gamma \in ISO(2, 1)$ are factorized into

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New particle limit

$$\gamma = \mathfrak{j}\,\mathfrak{p} = (\iota_3 + \iota^\mu J_\mu)(1 + \xi^\nu P_\nu)\,,\tag{14}$$

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with $j \in SO(2, 1)$, $\xi \equiv \xi^a P_a \in \mathfrak{so}(2, 1)^*$, and the group multiplication

$$\gamma_{(1)}\gamma_{(2)} = \mathfrak{j}_{(1)}\mathfrak{j}_{(2)}\left(1 + \mathrm{Ad}(\mathfrak{j}_{(2)}^{-1})\xi_{(1)} + \xi_{(2)}\right) \,. \tag{15}$$

The scalar product again has the form

$$\langle J_{\mu} P_{\nu} \rangle = \eta_{\mu\nu} , \qquad \langle J_{\mu} J_{\nu} \rangle = \langle P_{\mu} P_{\nu} \rangle = 0 .$$
 (16)

We can now put the Lagrangian in the boundary form

Particle in 3d gravity

$$L = \frac{k}{2\pi} \int_{\Gamma} \left\langle j^{-1} \dot{j} d\xi - \bar{j}^{-1} \dot{\bar{j}} d\bar{\xi} + \frac{C_J}{k} d\phi \left[\bar{j}^{-1} \dot{\bar{j}}, \bar{\xi} \right] + \frac{C_P}{k} d\phi \bar{j}^{-1} \dot{\bar{j}} \right\rangle, \quad (17)$$

where $C_J \equiv m J_0$, $C_P \equiv s P_0 = s S_0$. The sewing condition splits into

$$j^{-1} = \mathfrak{n} \, \boldsymbol{e}^{\frac{1}{k} \mathcal{C}_J \phi} \bar{j}^{-1} \,, \qquad -\mathrm{Ad}(\mathfrak{n}^{-1}) \, \boldsymbol{\xi} = \nu - \mathrm{Ad}(\boldsymbol{e}^{\frac{1}{k} \mathcal{C}_J \phi}) \, \bar{\boldsymbol{\xi}} + \frac{1}{k} \mathcal{C}_P \phi \,, \quad (18)$$

where $N = (1 + \nu) \mathfrak{n}$, $\mathfrak{n} \in SO(2, 1)$, $\nu \in \mathfrak{so}(2, 1)^*$. Substituting these conditions and denoting $\kappa \equiv \frac{k}{2\pi}$ we eventually obtain the Lagrangian

$$L = \kappa \left(\dot{\Pi}^{-1} \Pi \right)_{\mu} x^{\mu} + s \left(\mathfrak{n}^{-1} \dot{\mathfrak{n}} \right)_{0} , \qquad (19)$$

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with the particle's position $x \equiv \mathfrak{n} \overline{\xi} \mathfrak{n}^{-1} \in \mathfrak{so}(2,1)^*$ and momentum $\Pi \equiv \mathfrak{n} e^{\frac{m}{\kappa} J_0} \mathfrak{n}^{-1} \in \mathrm{SO}(2,1).$

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Properties of the gravitating particle

The parallel transport around the particle is described by the holonomy of the connection A_S along the boundary Γ , which is given by

In particular, $j(\phi = 0)j^{-1}(\phi = 2\pi) = \Pi$. The momentum manifold SO(2, 1) is 3d anti-de Sitter space. Using the parametrization $\Pi = p_3 + \frac{1}{\kappa}p^{\mu}J_{\mu}$, where $p_3 = \sqrt{1 - \frac{1}{4\kappa^2}p_{\mu}p^{\mu}}$, we find the mass shell condition

$$p_{\mu}p^{\mu} = 4\kappa^2 \sin^2 \frac{m}{2\kappa} \,. \tag{21}$$

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Properties of the gravitating particle – cont.

Let us restrict to the spinless case. With $x = x^{\mu}P_{\mu}$ and the Lagrange multiplier λ we may rewrite the effective Lagrangian as

$$L = -\left(p_{3}\dot{p}_{\mu} - \dot{p}_{3}p_{\mu} - \frac{1}{2\kappa}\epsilon_{\mu\nu\sigma}\dot{p}^{\nu}p^{\sigma}\right)x^{\mu} - \lambda\left(p_{\mu}p^{\mu} - 4\kappa^{2}\sin^{2}\frac{m}{2\kappa}\right).$$
(22)

It still leads to the equations of motion of a free relativistic particle

$$\dot{x}^{\mu} = 2\lambda \, \rho^{\mu} \,, \qquad \dot{\rho}_{\mu} = 0$$
 (23)

and in the limit $\kappa \to \infty$ we recover the free particle Lagrangian

$$L = -\dot{p}_{\mu}x^{\mu} - \lambda \left(p_{\mu}p^{\mu} - m^{2}\right) .$$
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The case of multiple particles

The Chern-Simons Lagrangian for a system of *n* particles has the form

$$L_{(n)} = \frac{k}{4\pi} \int_{S} \left\langle \dot{A}_{S} \wedge A_{S} \right\rangle - \sum_{i=1}^{n} \left\langle C_{i} h_{i}^{-1} \dot{h}_{i} \right\rangle + \int_{S} \left\langle A_{t} \left(\frac{k}{2\pi} F_{S} - \sum_{i=1}^{n} h_{i} C_{i} h_{i}^{-1} \delta^{2} (\vec{y} - \vec{y}_{i}) \, dy^{1} \wedge dy^{2} \right) \right\rangle.$$
(25)

We may divide S into *n* particle discs D_i and the empty polygon \mathcal{E} , whose edges Γ_i coincide with the boundaries of D_i 's. On each D_i the connection is given by

$$\boldsymbol{A}_{\mathcal{S}}^{(\mathcal{D}_{i})} = \bar{\gamma}_{i} \frac{1}{k} \mathcal{C}_{i} \boldsymbol{d} \phi_{i} \bar{\gamma}_{i}^{-1} + \bar{\gamma}_{i} \boldsymbol{d} \bar{\gamma}_{i}^{-1}, \quad \bar{\gamma}_{i} (\boldsymbol{r}_{i} = \boldsymbol{0}) = \boldsymbol{h}_{i}.$$
(26)

Then for the i'th particle we can derive

$$L_{i} = \kappa \left(\dot{\Pi}_{i}^{-1} \Pi_{i} \right)_{\mu} \boldsymbol{x}_{i}^{\mu} + \boldsymbol{s}_{i} \left(\boldsymbol{\mathfrak{n}}_{i}^{-1} \dot{\boldsymbol{\mathfrak{n}}}_{i} \right)_{\boldsymbol{0}} \,. \tag{27}$$

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The case of multiple particles – cont.

Imposing the continuity at the polygon's vertices, $\gamma(\phi_{i+1} = 0) = \gamma(\phi_i = 2\pi)$ and fixing $\gamma(\phi_1 = 0) = 1$ we obtain the sequence of conditions

$$\mathfrak{n}_1 \overline{\mathfrak{j}}_1^{-1} = 1, \qquad \mathfrak{n}_2 \overline{\mathfrak{j}}_2^{-1} = \overline{\Pi}_1, \qquad \mathfrak{n}_3 \overline{\mathfrak{j}}_3^{-1} = \overline{\Pi}_1 \overline{\Pi}_2, \qquad \dots,$$
 (28)

where $\bar{\Pi}_i = \bar{\mathfrak{j}}_i e^{\frac{1}{\kappa}m_i J_0} \bar{\mathfrak{j}}_j^{-1}$. Using them we find the *n*-particle Lagrangian

$$L_{(n)} = \sum_{i=1}^{n} \left(\kappa \left(\dot{\Pi}_{i}^{-1} \bar{\Pi}_{i} \right)_{\mu} \bar{x}_{i}^{\mu} + s_{i} \left(\bar{j}_{i}^{-1} \dot{\bar{j}}_{i} \right)_{0} \right) + \kappa \left(\bar{\Pi}_{2}^{-1} \dot{\Pi}_{1}^{-1} \bar{\Pi}_{1} \bar{\Pi}_{2} - \dot{\Pi}_{1}^{-1} \bar{\Pi}_{1} \right)_{\mu} \bar{x}_{2}^{\mu} + s_{2} \left(\bar{j}_{2}^{-1} \bar{\Pi}_{1}^{-1} \dot{\Pi}_{1} \bar{j}_{2} \right)_{0} + \dots, \quad (29)$$

where $\bar{x}_i = \bar{\mathfrak{j}}_i \bar{\xi}_i \bar{\mathfrak{j}}_i^{-1}$.

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Alternative contraction of SO(3, 1)

Let us rescale the $\mathfrak{so}(3,1)$ generators to $\tilde{J}_{\mu} \equiv \sqrt{\Lambda} J_{\mu}$, $\tilde{P}_{\mu} \equiv 1/\sqrt{\Lambda} P_{\mu}$ and define $\tilde{S}_{\mu} \equiv \tilde{P}_{\mu} + \epsilon_{\mu 0\nu} \tilde{J}^{\nu}$. In the limit $\Lambda \to 0$ we obtain the algebra

$$\begin{split} & [\tilde{J}_{\mu}, \tilde{J}_{\nu}] = 0, \qquad [\tilde{J}_{\mu}, \tilde{S}_{\nu}] = \eta_{\nu 0} \tilde{J}_{\mu} - \eta_{\mu \nu} \tilde{J}_{0}, \\ & [\tilde{S}_{\mu}, \tilde{S}_{\nu}] = \eta_{\mu 0} \tilde{S}_{\nu} - \eta_{\nu 0} \tilde{S}_{\mu}. \end{split}$$
(30)

It generates the group $AN(2) \ltimes \mathfrak{an}(2)^*$, whose elements are

$$\gamma = \mathfrak{i}\mathfrak{s} = (1 + \iota^{\mu}\tilde{J}_{\mu})(\xi_{3} + \xi^{\nu}\tilde{S}_{\nu}), \qquad (31)$$

with $\mathfrak{s} \in AN(2)$, $\iota \equiv \iota^{\mu} \tilde{J}_{\mu} \in \mathfrak{an}(2)^* \simeq \mathbb{R}^3$, and the group multiplication

$$\gamma_{(1)}\gamma_{(2)} = \left(1 + \iota_{(1)} + \mathrm{Ad}(\mathfrak{s}_{(1)})\,\iota_{(2)}\right)\mathfrak{s}_{(1)}\mathfrak{s}_{(2)}\,. \tag{32}$$

The corresponding scalar product is

$$\left\langle \tilde{J}_{\mu}\tilde{S}_{\nu}\right\rangle = \eta_{\mu\nu}, \qquad \left\langle \tilde{J}_{\mu}\tilde{J}_{\nu}\right\rangle = \left\langle \tilde{S}_{\mu}\tilde{S}_{\nu}\right\rangle = 0.$$
 (33)

New effective particle Lagrangian

We also have to exchange the labels of mass and spin so that $C_{\tilde{J}} = s \tilde{J}_0$ and $C_{\tilde{S}} = m \tilde{S}_0$. Then we put the Lagrangian in the boundary form

$$L = \frac{k}{2\pi} \int_{\Gamma} \left\langle \dot{\Im}\Im^{-1} \left(d\bar{\mathfrak{s}}\,\bar{\mathfrak{s}}^{-1} - \bar{\mathfrak{s}}\,\frac{1}{k}\mathcal{C}d\phi\,\bar{\mathfrak{s}}^{-1} \right) + \frac{1}{k}\mathcal{C}d\phi\,\bar{\mathfrak{s}}^{-1}\dot{\bar{\mathfrak{s}}} \right\rangle\,,\qquad(34)$$

where we denote $\mathfrak{I}\equiv\overline{\mathfrak{j}}^{-1}\mathfrak{j}.$ The sewing condition can be split into

$$\mathfrak{s}^{-1} = \mathfrak{v} \, e^{\frac{1}{k} \, \mathcal{C}_{\bar{S}} \phi} \bar{\mathfrak{s}}^{-1} \,, \qquad \mathfrak{I} = e^{-\frac{1}{k} \mathcal{C}_{\bar{J}} \phi} \mathfrak{s} \, (1 - n) \, \mathfrak{s}^{-1} \,, \tag{35}$$

where N = (1 + n) v, $v \in AN(2)$, $n \in an(2)^*$. Substituting it we obtain the final Lagrangian

$$L = \kappa \left(\dot{\Pi} \Pi^{-1} \right)_{\mu} x^{\mu} + s \left(\bar{\mathfrak{s}}^{-1} \dot{\bar{\mathfrak{s}}} \right)_{0} , \qquad (36)$$

with the particle's momentum $\Pi \equiv \bar{\mathfrak{s}} e^{\frac{m}{\kappa}\bar{S}_0}\bar{\mathfrak{s}}^{-1} \in AN(2)$ and position $x \equiv \bar{\mathfrak{s}} \mathfrak{v}^{-1} n \mathfrak{v} \bar{\mathfrak{s}}^{-1} \in \mathfrak{an}(2)^*$.

Properties of the particle

The holonomy of the connection A_S along the boundary Γ is given by

$$\operatorname{hol}_{\Gamma}(A_{S}) = \gamma(\phi = 0) \gamma^{-1}(\phi = 2\pi) = (1 + (1 - \operatorname{Ad}(\Pi)) x + \frac{1}{\kappa} C_{\tilde{J}}) \Pi.$$
(37)

In particular, $\mathfrak{s}(\phi = 0)\mathfrak{s}^{-1}(\phi = 2\pi) = \Pi$. The momentum manifold AN(2) is 3d de Sitter space. Furthermore, using the parametrization $\Pi = e^{p^a/\kappa} \tilde{S}_a e^{p^0/\kappa} \tilde{S}_0$, $\mathfrak{v} = e^{\upsilon^a/\kappa} \tilde{S}_a e^{\upsilon^0/\kappa} \tilde{S}_0$, a = 1, 2 we find that

$$p^0 = m, \qquad p^a = \left(1 - e^{\frac{m}{\kappa}}\right) v^a,$$
 (38)

i.e. the particle's energy is fixed to be the rest energy.

Properties of the particle – cont.

The spin term does not contribute to the equations of motion and we may restrict to s = 0. With $x = x^{\mu}\tilde{P}_{\mu}$ and the Lagrange multiplier λ the particle Lagrangian can be rewritten in the form

$$L = -(x^{0}\dot{p}_{0} + x^{a}\dot{p}_{a} - \kappa^{-1}x^{a}p_{a}\dot{p}_{0}) - \lambda(p_{0}^{2} - m^{2}), \quad (39)$$

which describes a κ -deformed Carroll particle. It gives the equations of motion of an ordinary Carroll particle

$$\dot{x}^0 = 2\lambda m, \qquad \dot{x}^a = 0, \qquad \dot{p}_\mu = 0$$
 (40)

and in the limit $\kappa \to \infty$ we recover the Carroll particle Lagrangian

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$$L = -x^{0}\dot{p}_{0} - x^{a}\dot{p}_{a} - \lambda \left(p_{0}^{2} - m^{2}\right) . \tag{41}$$

Symmetries of the deformed Carroll particle

The particle is invariant under infinitesimal κ -deformed Carroll transformations, which include ordinary rotations

$$\delta x^{a} = \rho \,\epsilon^{a}_{\ b} x^{b} \,, \qquad \delta p_{a} = \rho \,\epsilon^{\ b}_{a} p_{b} \,, \qquad \delta x^{0} = \delta p_{0} = 0 \,, \qquad (42)$$

deformed Carrollian boosts

$$\delta x^{0} = \left(1 + \kappa^{-1} p_{0}\right) \lambda_{a} x^{a}, \qquad \delta p_{a} = -\lambda_{a} p_{0}, \qquad \delta x^{a} = \delta p_{0} = 0,$$
 (43)

deformed translations

$$\delta x^{0} = \alpha^{0}, \qquad \delta x^{a} = e^{p_{0}/\kappa} \alpha^{a}, \qquad \delta p_{\mu} = 0$$
(44)

and spatial conformal transformations

$$\delta x^{a} = \eta x^{a}, \qquad \delta p_{a} = -\eta p_{a}, \qquad \delta x^{0} = \delta p_{0} = 0, \qquad (45)$$

where ρ , λ_a , α^{μ} , η denote transformation parameters.

Multiple particles

At the boundary Γ_i of each disc \mathcal{D}_i , $i = 1, \ldots, n$ we find

$$L = \kappa \left(\dot{\Pi}_i \, \Pi_i^{-1} \right)_{\mu} x_i^{\mu} + s_i \left(\bar{\mathfrak{s}}_i^{-1} \dot{\bar{\mathfrak{s}}}_i \right)_0 \,. \tag{46}$$

The continuity $\gamma(\phi_{i+1} = 0) = \gamma(\phi_i = 2\pi)$ and fixing $\gamma(\phi_1 = 0) = 1$ leads to the sequence of conditions

$$\mathfrak{v}_1\bar{\mathfrak{s}}_1^{-1}=1\,,\qquad\mathfrak{v}_2\bar{\mathfrak{s}}_2^{-1}=\bar{\Pi}_1\,,\qquad\mathfrak{v}_3\bar{\mathfrak{s}}_3^{-1}=\bar{\Pi}_2\bar{\Pi}_1\,,\qquad\ldots\,,\qquad(47)$$

where $\bar{\Pi}_i = v_i e^{\frac{1}{\kappa} m_i \tilde{S}_0} v_i^{-1}$. Then we obtain the *n*-particle Lagrangian

$$L_{(n)} = \sum_{i=1}^{n} \left(\kappa \left(\dot{\bar{\Pi}}_{i} \bar{\Pi}_{i}^{-1} \right)_{\mu} \bar{x}_{i}^{\mu} + s_{i} \left(v_{i}^{-1} \dot{v}_{i} \right)_{0} \right) + \\ \kappa \left(\bar{\Pi}_{2} \dot{\bar{\Pi}}_{1} \bar{\Pi}_{1}^{-1} \bar{\Pi}_{2}^{-1} - \dot{\bar{\Pi}}_{1} \bar{\Pi}_{1}^{-1} \right)_{\mu} \bar{x}_{2}^{\mu} + s_{2} \left(v_{2}^{-1} \bar{\Pi}_{1} \dot{\bar{\Pi}}_{1}^{-1} v_{2} \right)_{0} + \dots, \quad (48)$$

where $\bar{x}_i = n_i$.



- We rederived the action of a gravitating particle in flat spacetime
- We obtained the new action of a κ-deformed Carroll particle
- Its momentum space is equivalent to the 3d κ-Minkowski momentum space, associated with the κ-Poincaré Hopf algebra
- The relevance of this particle model remains to be understood

References

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